



# A TALE OF TWO UNCERTAINTIES:

*Analyzing  $P_T$  Bias and its Effects on the Dimuon Mass Spectrum*

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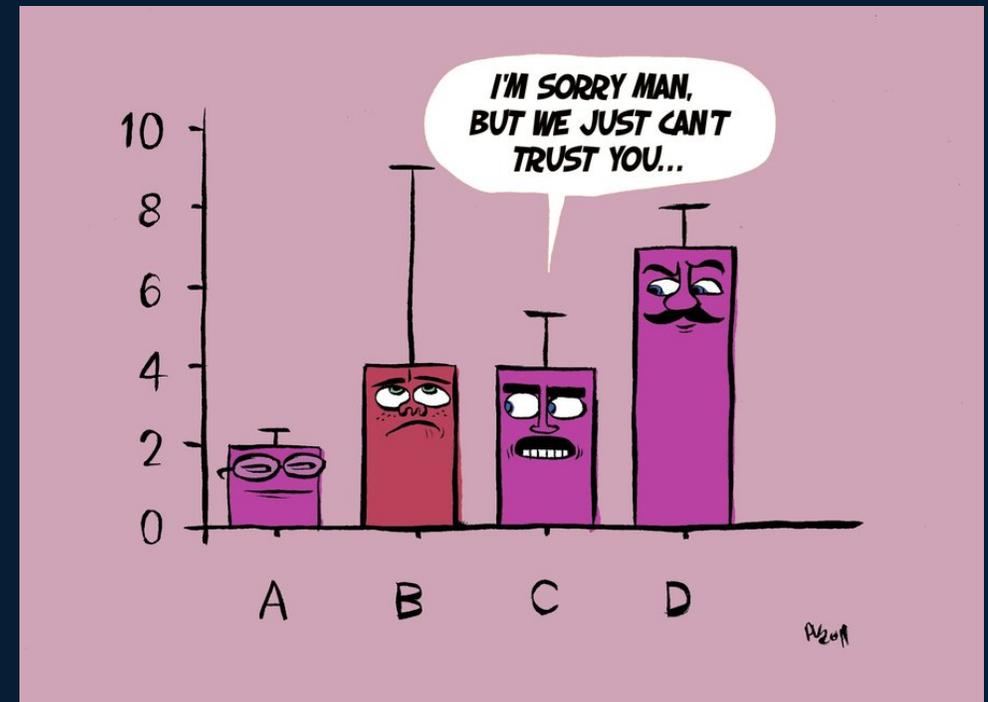
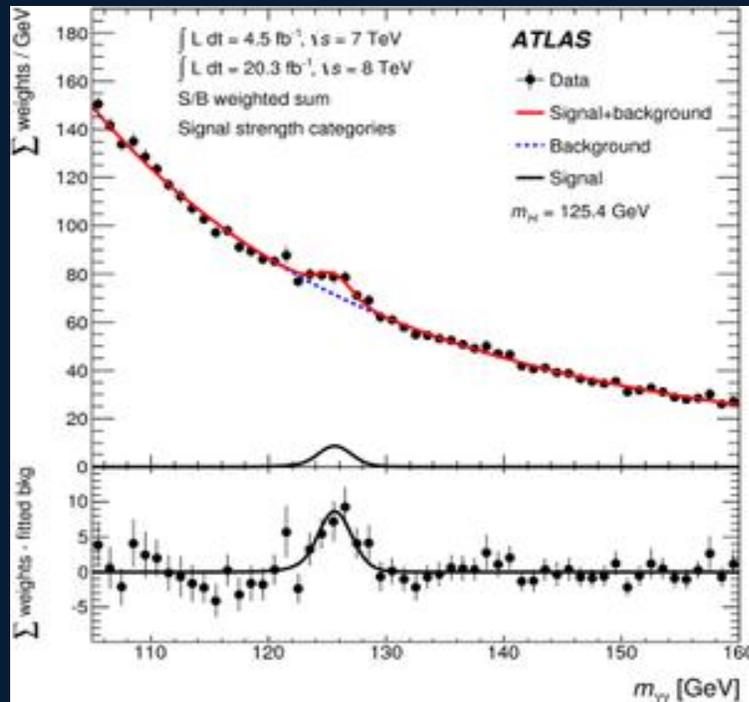
Tamra Nebabu, *Duke University*

Dr. Pushpalatha Bhat, Dr. Leonard Spiegel

Fermilab, CMS

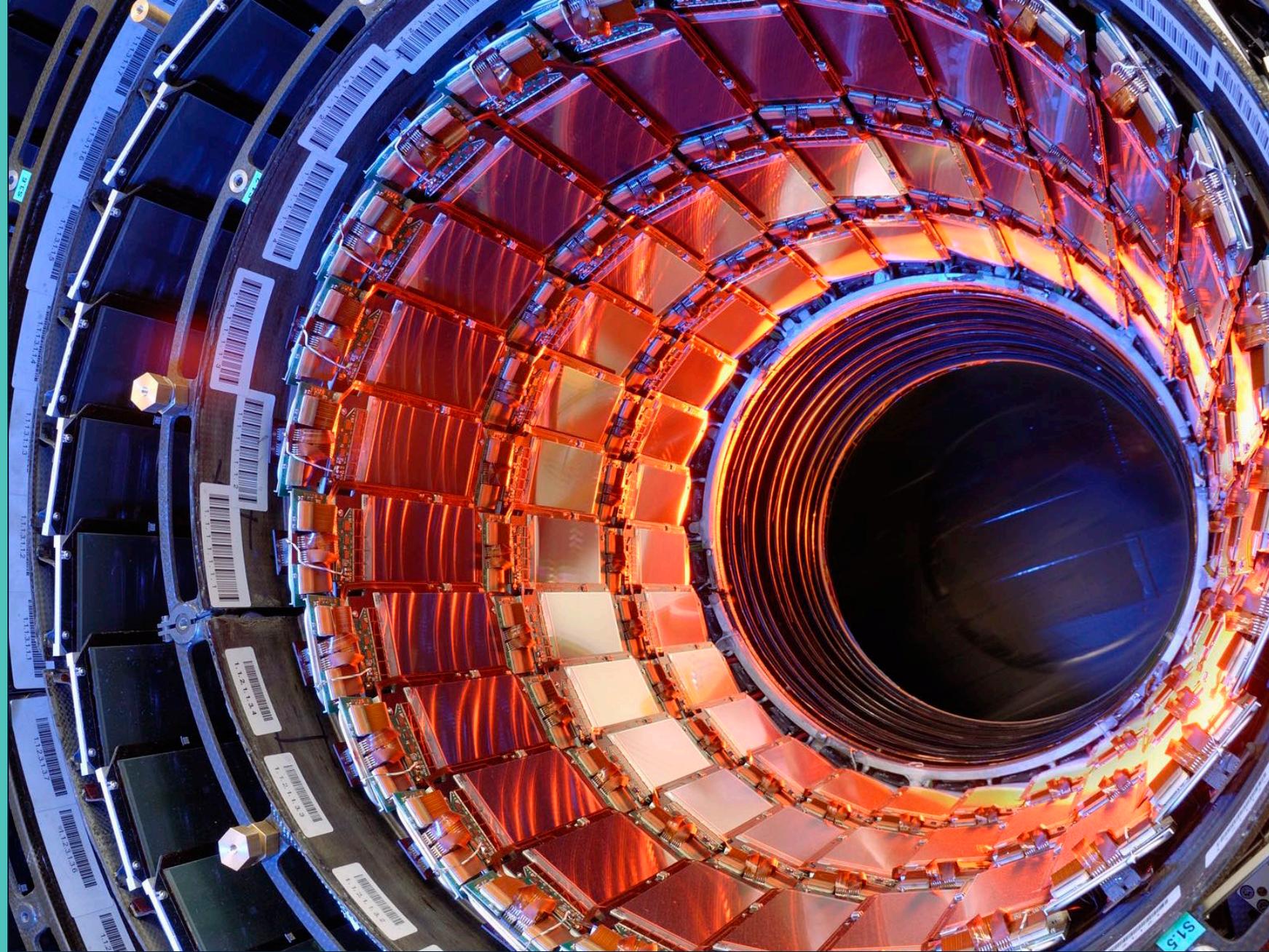
# WHY UNCERTAINTY MATTERS

- It changes conclusions!
  - Ex: particle identification based on peaks

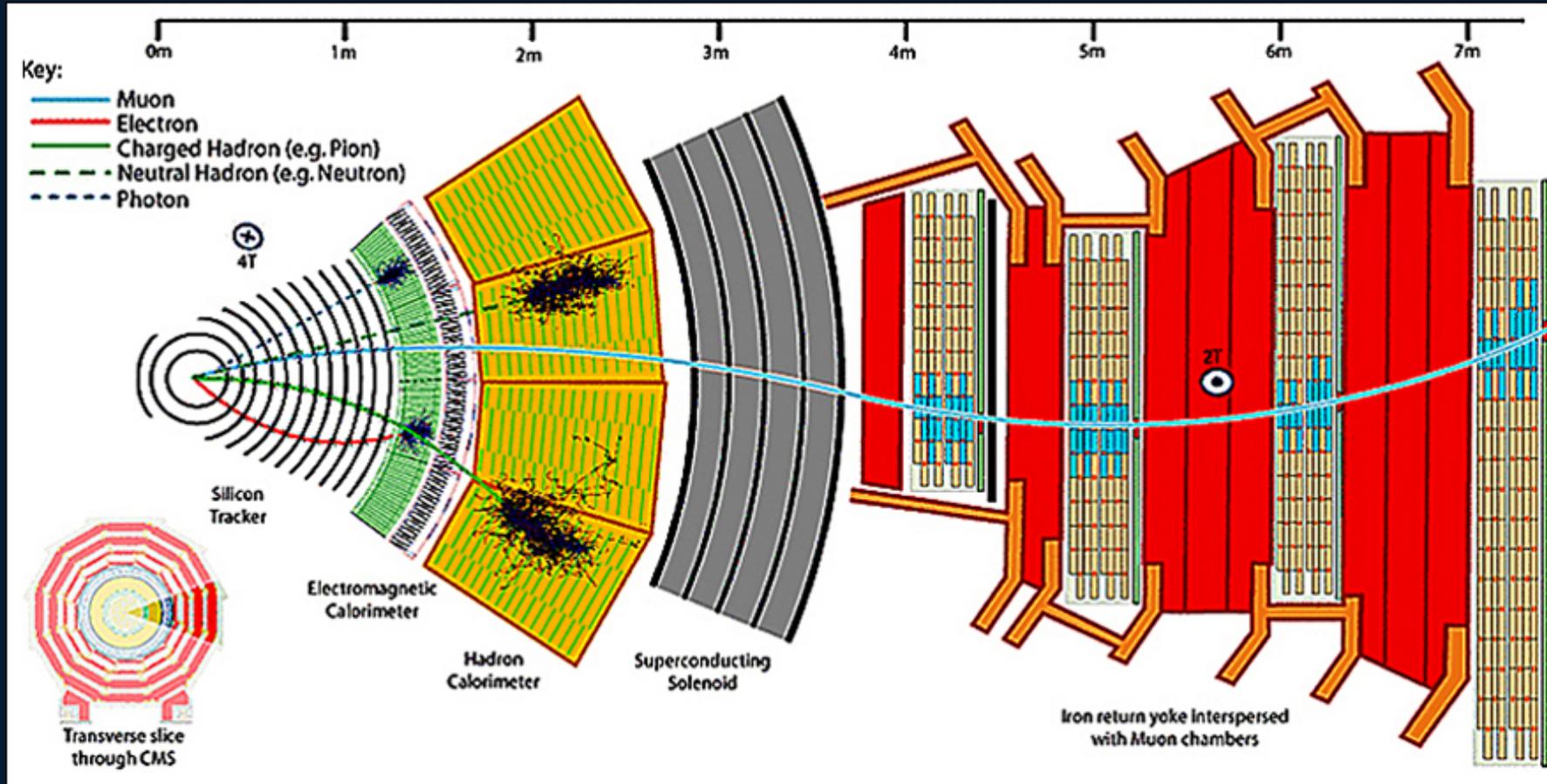


# COMPACT MUON SOLENOID (CMS) DETECTOR

A general purpose  
LHC detector



# THE LAYERS OF CMS



# CMS SPECS

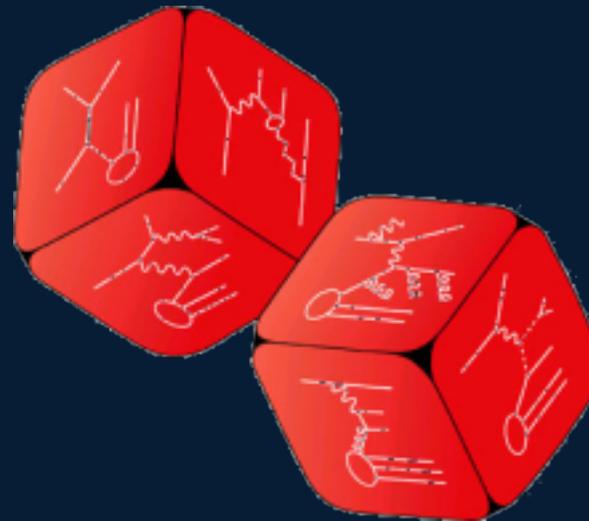
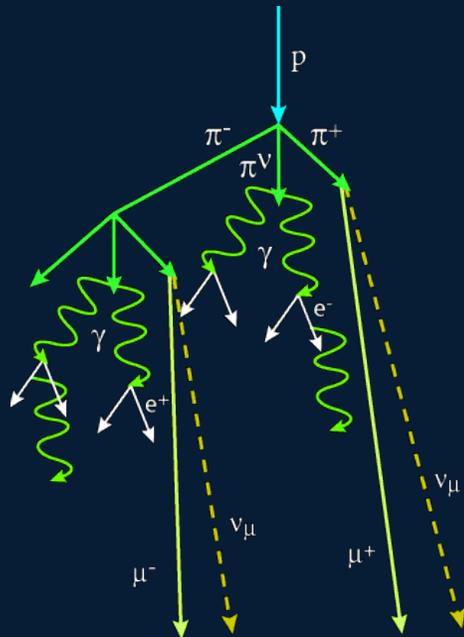
- Heaviest LHC detector
- 2<sup>nd</sup> largest general purpose detector in volume
- 100 meters underground
- 14 million kgs = 14,000  = 5,000 
- 4 T magnet = 100,000  = 2.7 
- 5 layers – silicon tracker, EM calorimeter, hadron calorimeter, solenoidal magnet, muon detection layer

# MY DATA

Cosmic muon data  
collected from CMS in  
2015

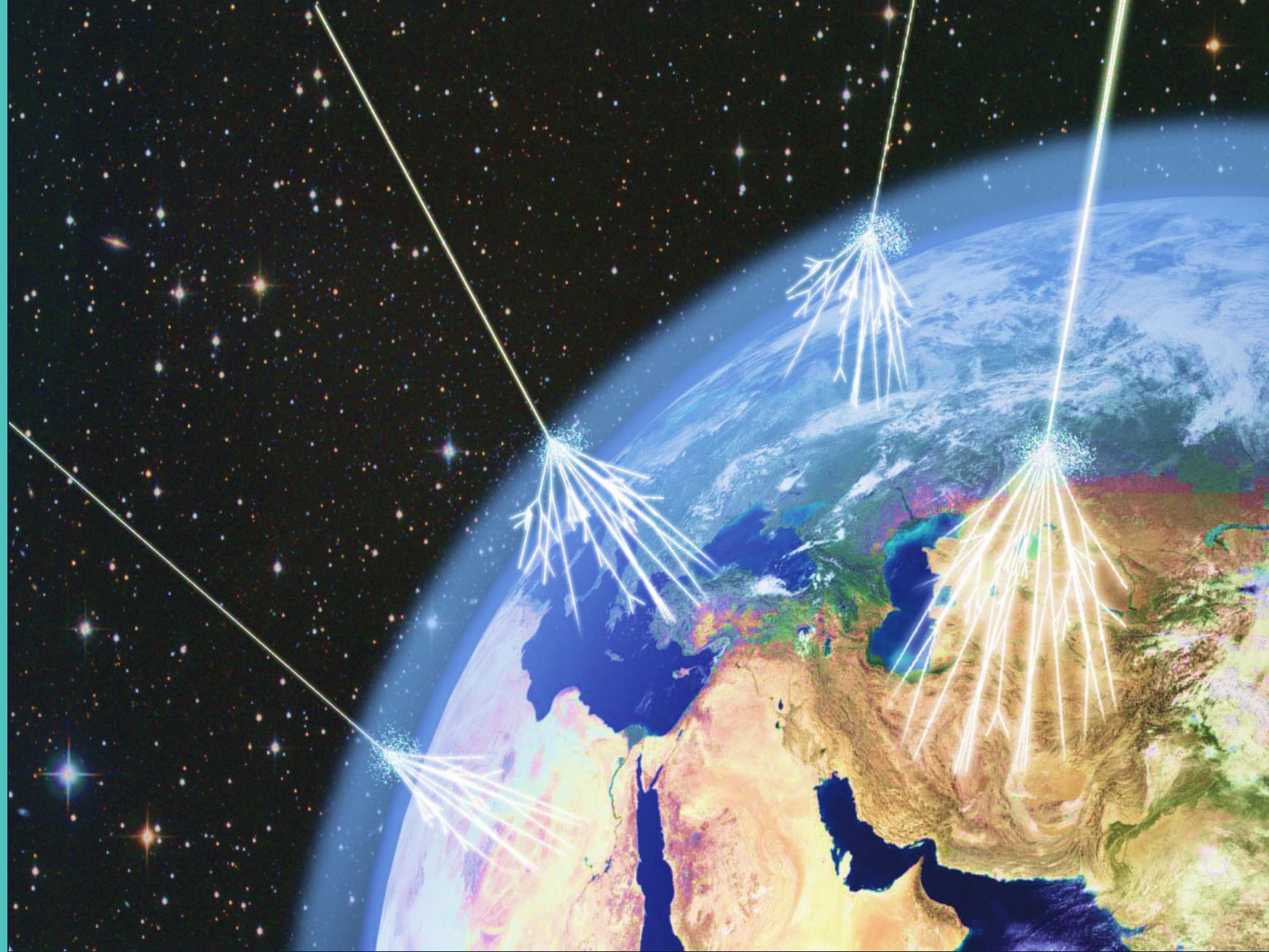
Monte carlo  
simulations of cosmic  
muons

Simulations of  
collision events  
generated in Pythia8

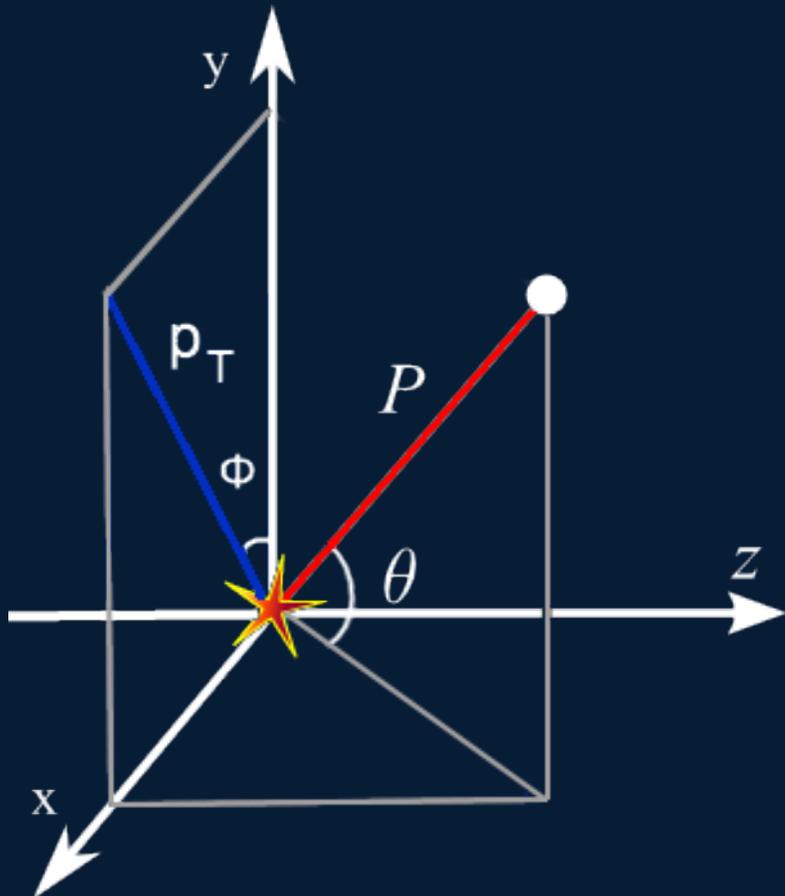


# COSMIC ENDPOINT METHOD

Measuring  
uncertainty in  $p_T$



# WHAT KIND OF MEASUREMENTS?



- These variables can be used to calculate virtually every property of the particle we're interested in!

- Physicists also like to use  $\eta$

$$\eta = -\ln\left(\tan\frac{\theta}{2}\right)$$

- In my study, I studied distributions of  $p_T$

# THE STUDY

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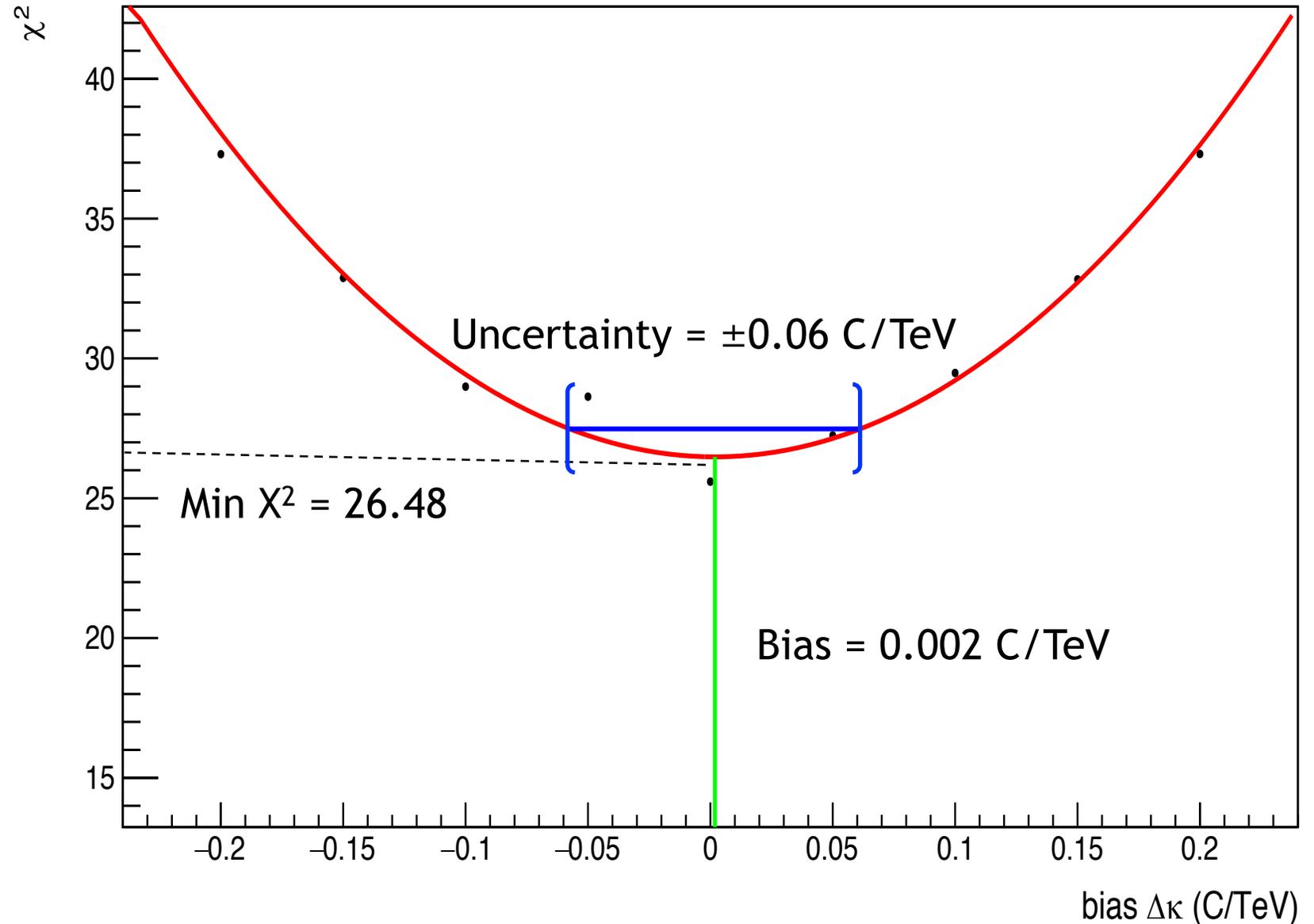
- Assume: **detector has systematic error or bias** that causes  $p_T$  to be scaled up by factor that depends on  $p_T$ 
  - $p_T' = \alpha p_T$  where  $\alpha$  depends linearly on  $p_T$
- This is equivalent to a **constant shift in curvature**  $\kappa = \frac{q}{p_T} = \frac{\pm 1}{p_T}$
- Procedure
  1. Make  $p_T$  histograms
  2. Calculate and make  $\kappa$  histograms
  3. Apply shift to  $\kappa$  of simulation
  4. Compare data and simulation
  5. Rinse and repeat

# $\chi^2$ vs. $\Delta\kappa$ (TRAD3) for 16 bins (CRAFT/ASYMP)

## RESULTS

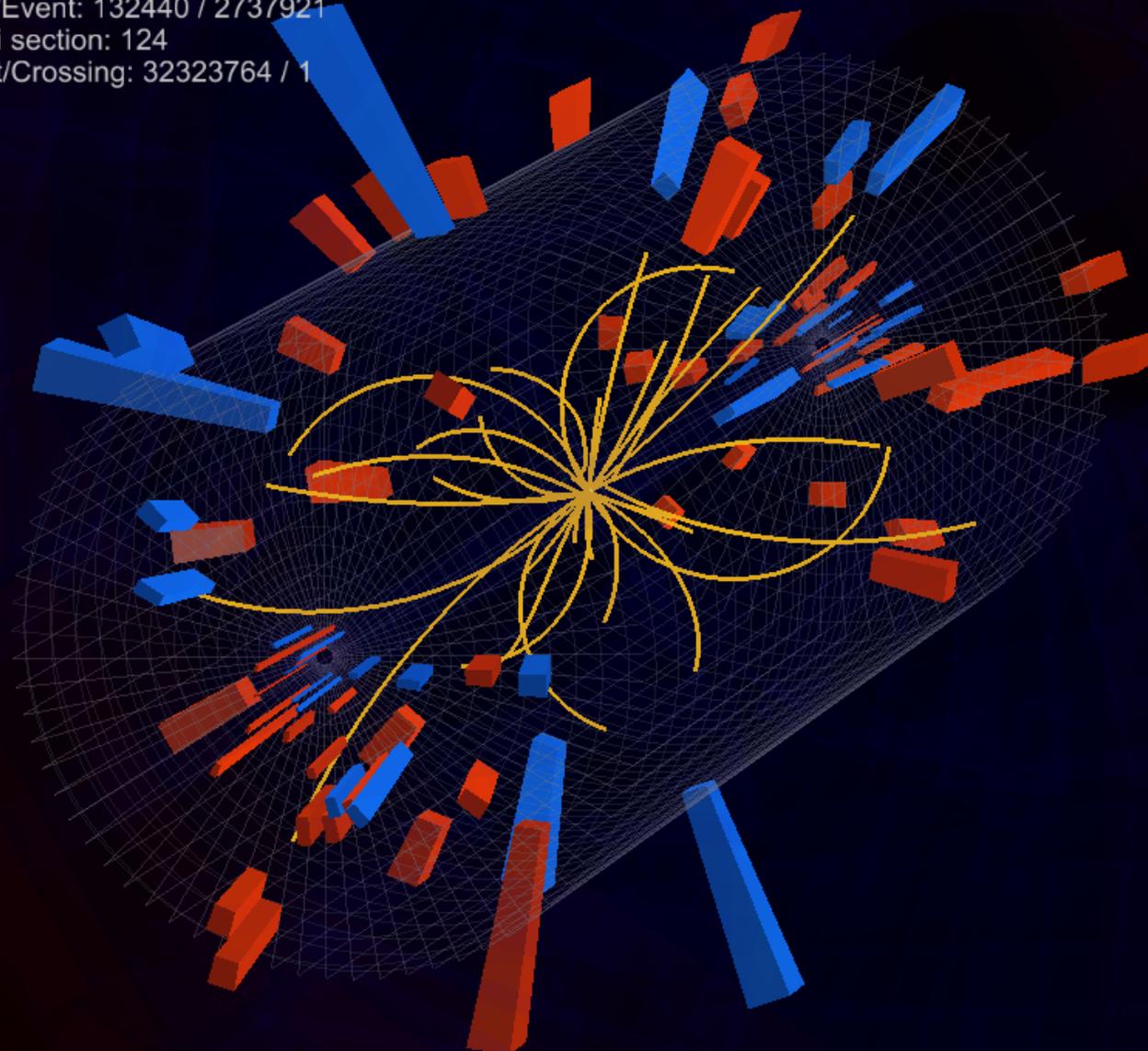
The bias in the detector is very small!

...but now we have uncertainty...





CMS Experiment at LHC, CERN  
Data recorded: Tue Mar 30 12:58:48 2010 CEST  
Run/Event: 132440 / 2737921  
Lumi section: 124  
Orbit/Crossing: 32323764 / 1

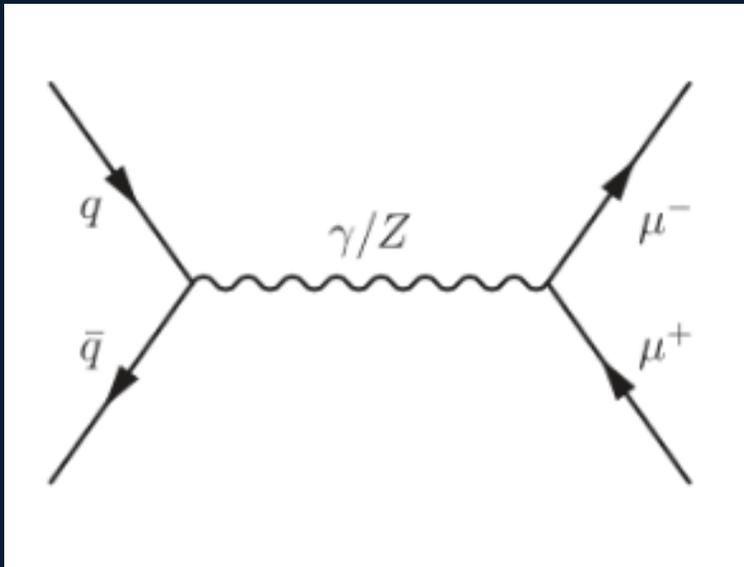


# RELATIVE DIFFERENCE STUDY

Propagating  
uncertainty in  $p_T$  to  
the mass spectrum

# QUARK COMPOSITENESS

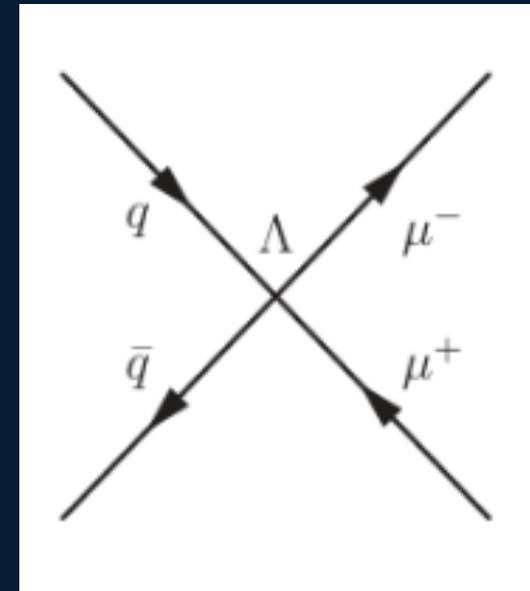
## Drell-Yan (DY)



$$q\bar{q} \rightarrow Z / \gamma^* \rightarrow l^+l^-$$

Z boson intermediate

## Contact Interaction (CI)



$$q\bar{q} \rightarrow l^+l^-$$

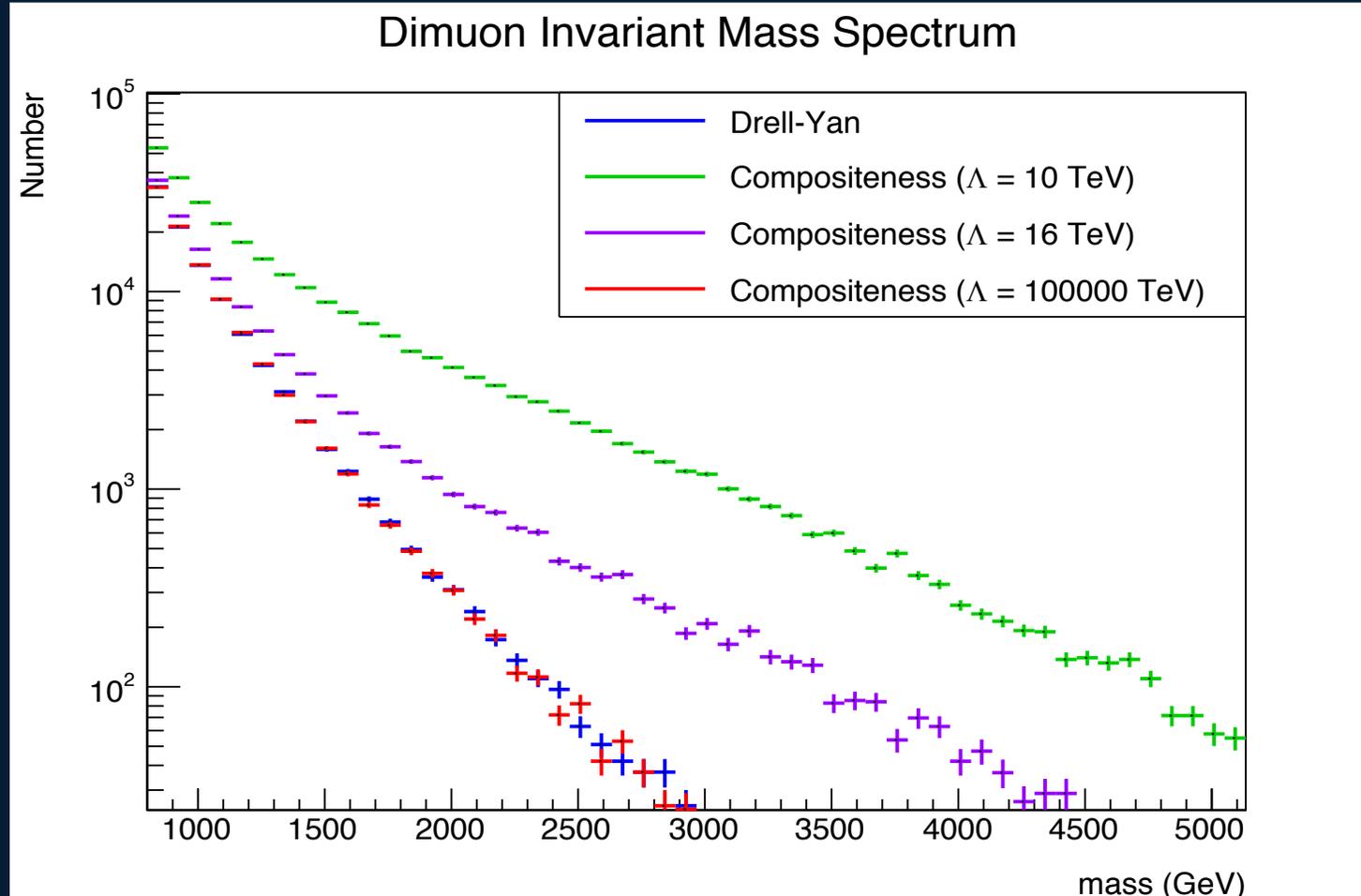
No Z boson intermediate

# INVARIANT MASS

- Is the **rest mass of the two-muon system**
  - **NOT the sum of the rest masses!!**
- By conservation of mass-energy, should be more or less equivalent to the mass of the parent Z boson
- Derived from  $E^2 = (m_0c^2)^2 + (pc^2)^2$
- In the highly relativistic ( $E \gg m$ ) case, approximates to

$$M = \sqrt{2p_{T_1}p_{T_2} (\cosh(\eta_1 - \eta_2) - \cos(\phi_1 - \phi_2))}$$

# DRELL-YAN VS. CONTACT INTERACTION



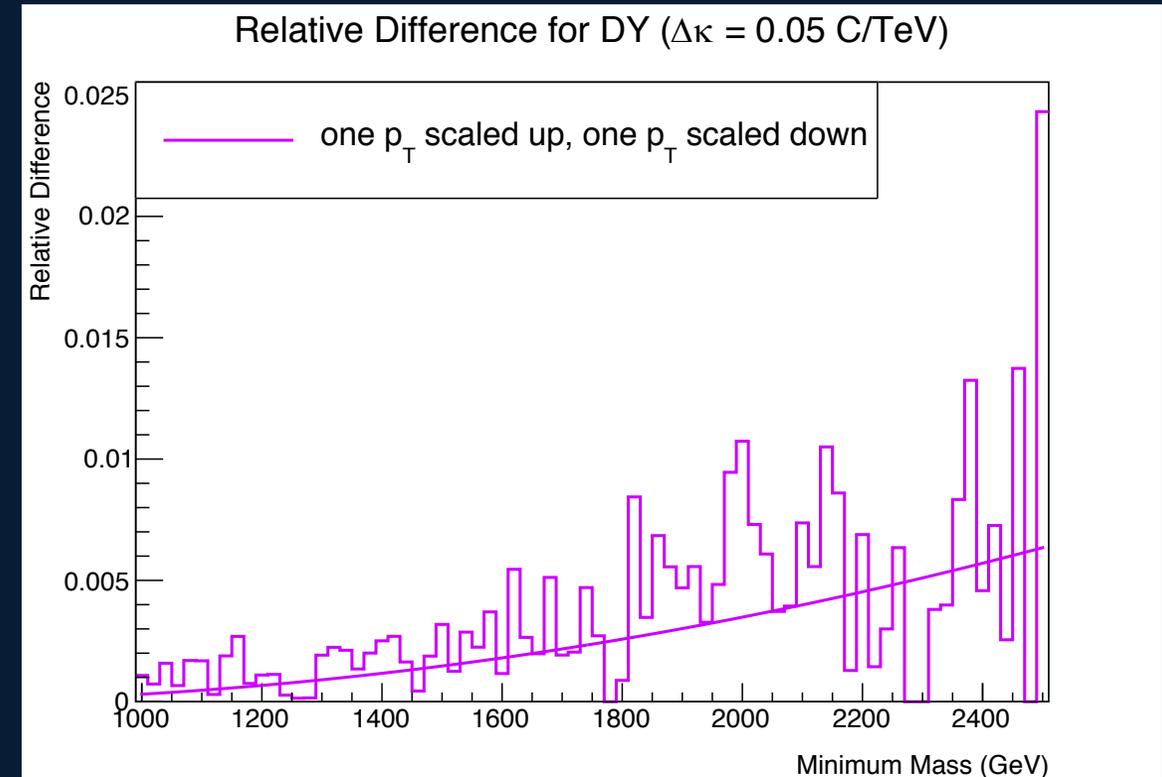
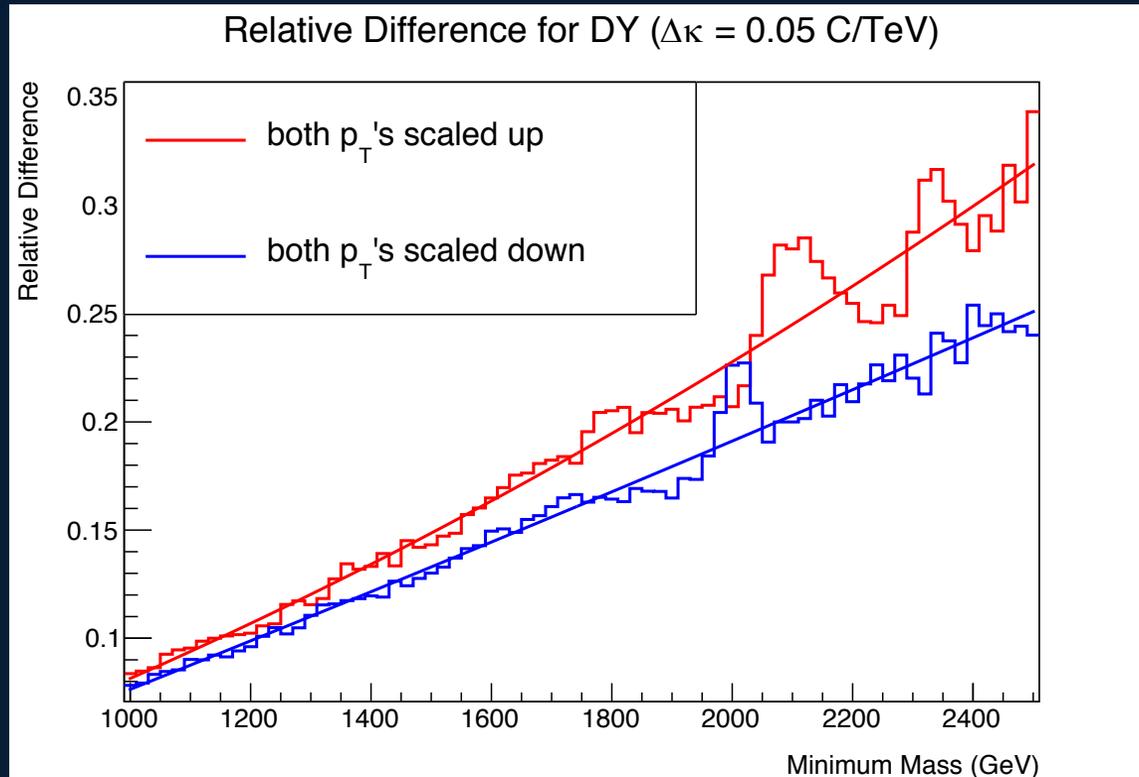
# THE STUDY

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- Recalculate the masses using a shifted value of  $p_T$ 
  - shifting  $\kappa \rightarrow$  shifted  $p_T \rightarrow$  shifted mass
- Perform a counting experiment!
  1. Pick a minimum mass
  2. Count up the number of entries above that mass value in the unshifted spectrum
  3. Count again for the shifted spectrum
  4. Calculate the relative difference between the shifted and unshifted mass spectra

$$\text{Relative Difference} = \frac{|\text{Shifted Integral} - \text{Unshifted Integral}|}{\text{Unshifted Integral}}$$

# RESULTS



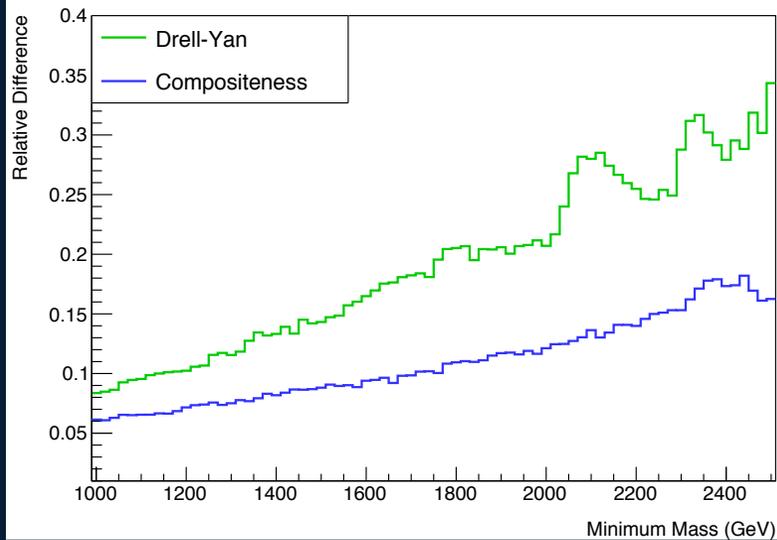
# DRELL-YAN VS. CONTACT INTERACTION

Scaling up

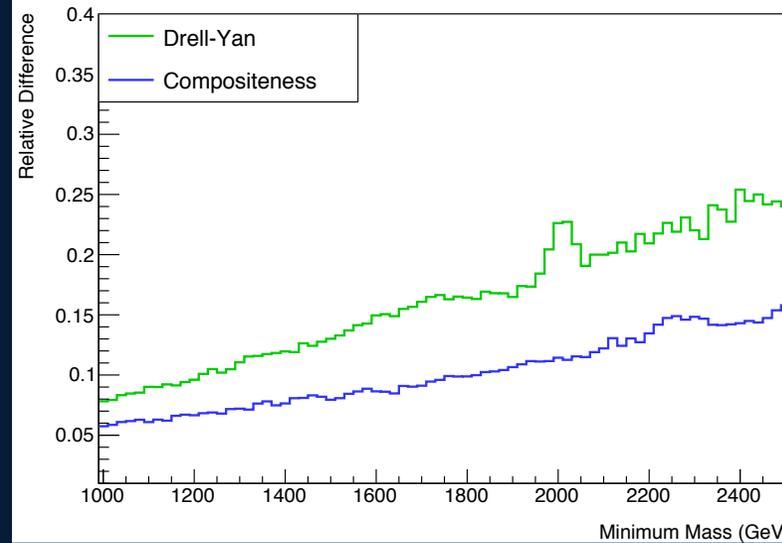
Scaling down

One up, one down

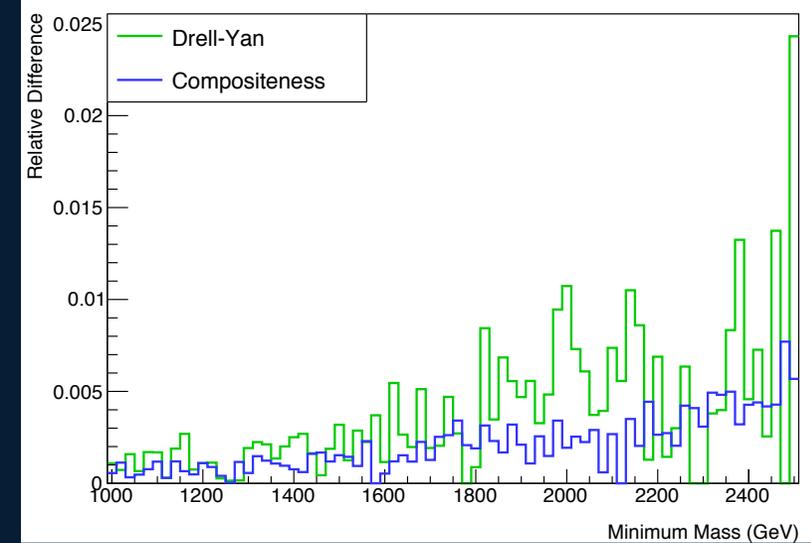
Relative Difference for DY vs. CI after Scaling Up ( $\Lambda = 16 \text{ TeV}$ ,  $\Delta\kappa = 0.05 \text{ C/TeV}$ )



Relative Difference for DY vs. CI after Scaling Down ( $\Lambda = 16 \text{ TeV}$ ,  $\Delta\kappa = 0.05 \text{ C/TeV}$ )

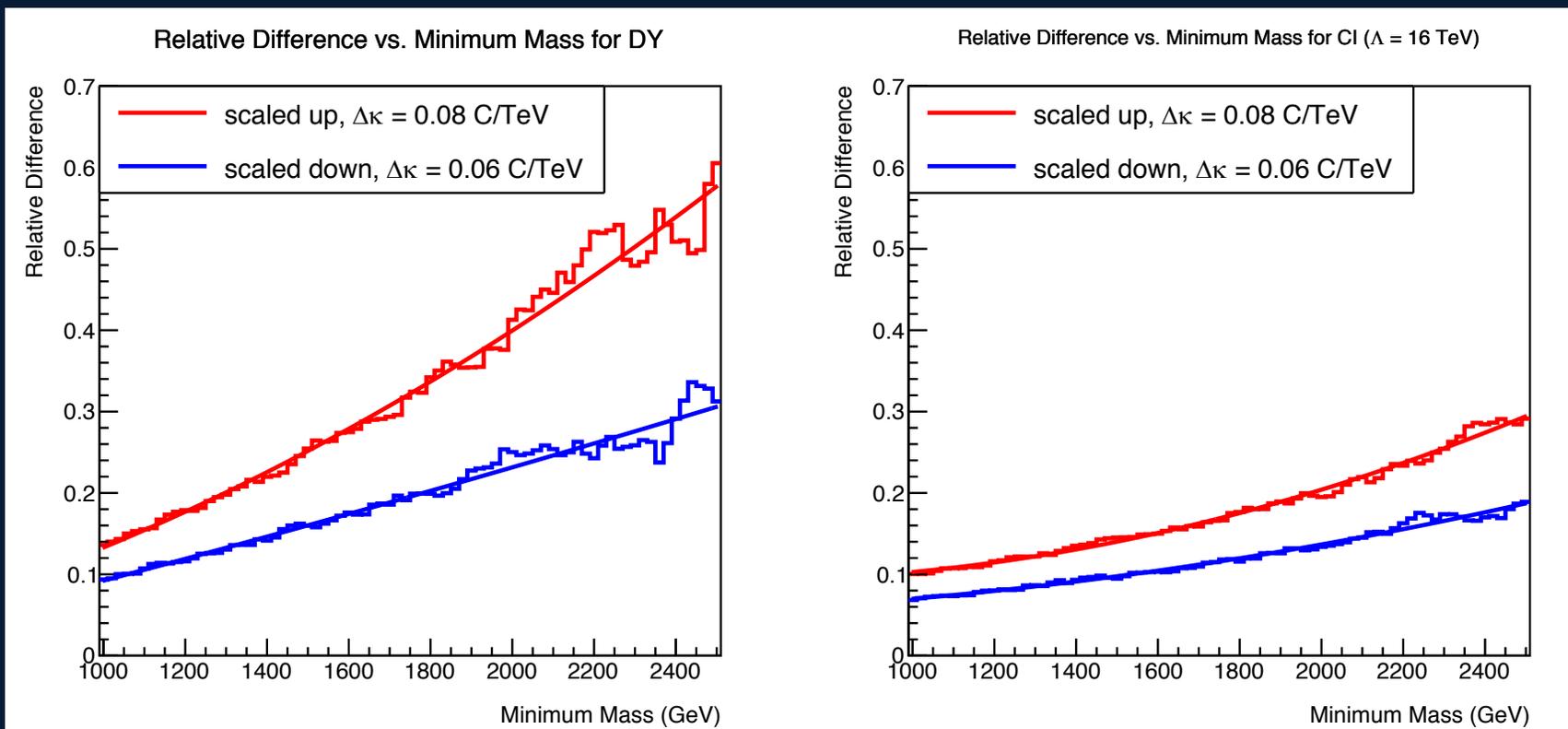


Relative Difference for DY vs. CI after Curvature Scaling ( $\Lambda = 16 \text{ TeV}$ ,  $\Delta\kappa = 0.05 \text{ C/TeV}$ )



# CONNECTING THE DOTS

Use the uncertainty in  $p_T$  obtained by the Cosmic Endpoint Method to get an **uncertainty band** in the dimuon mass spectrum



# ACKNOWLEDGEMENTS

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- My supervisors Dr. Pushpa Baht and Dr. Leonard Spiegel
- Shawn Zaleski
- My office buddy Amanda
- Dr. Elliott McCrory, Sandra Charles, and the entire SIST committee and staff
- My SIST mentors William Freeman and Jodi Coghill
- You guys!

I THOUGHT I WAS  
INTERESTED IN UNCERTAINTY  
BUT NOW I'M NOT SO SURE



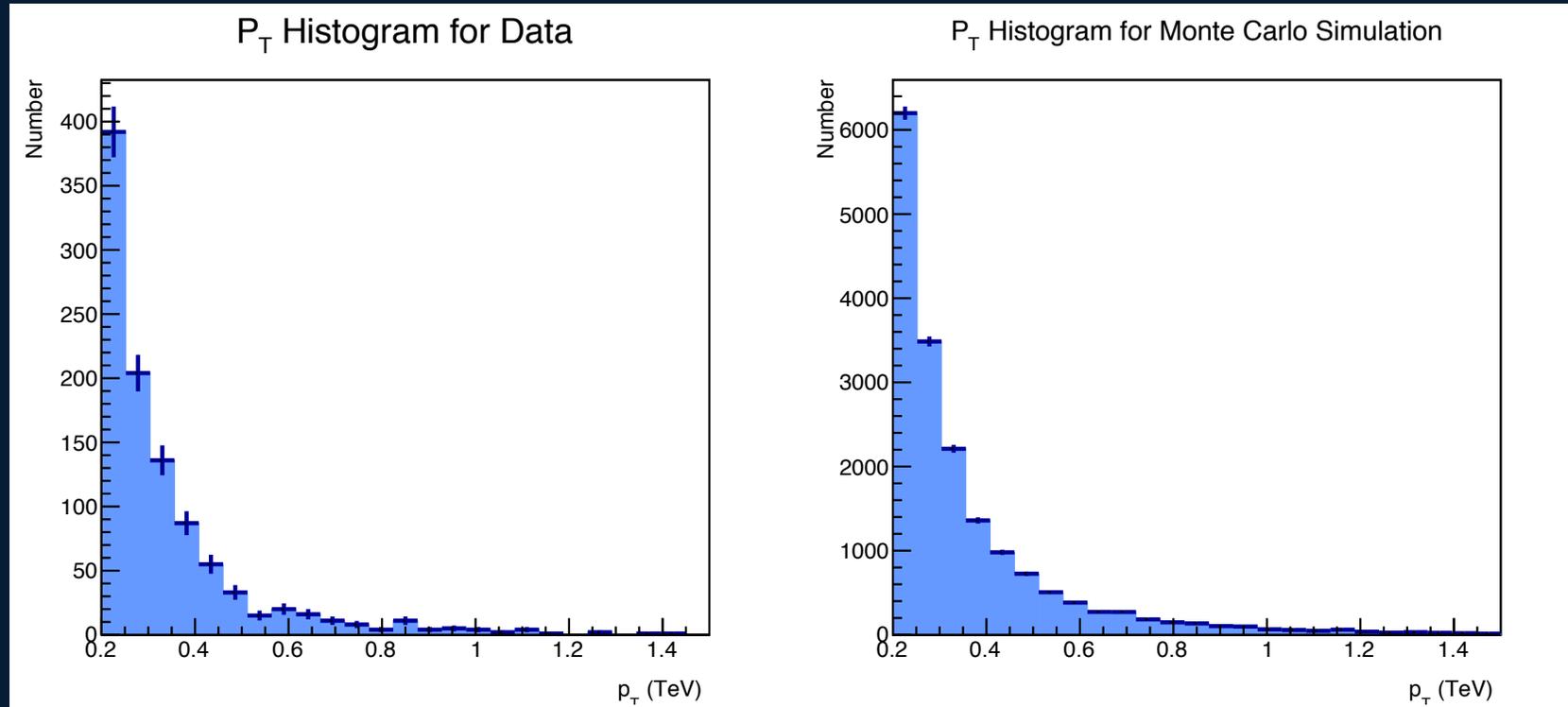
Josh's

# BACKUP SLIDES

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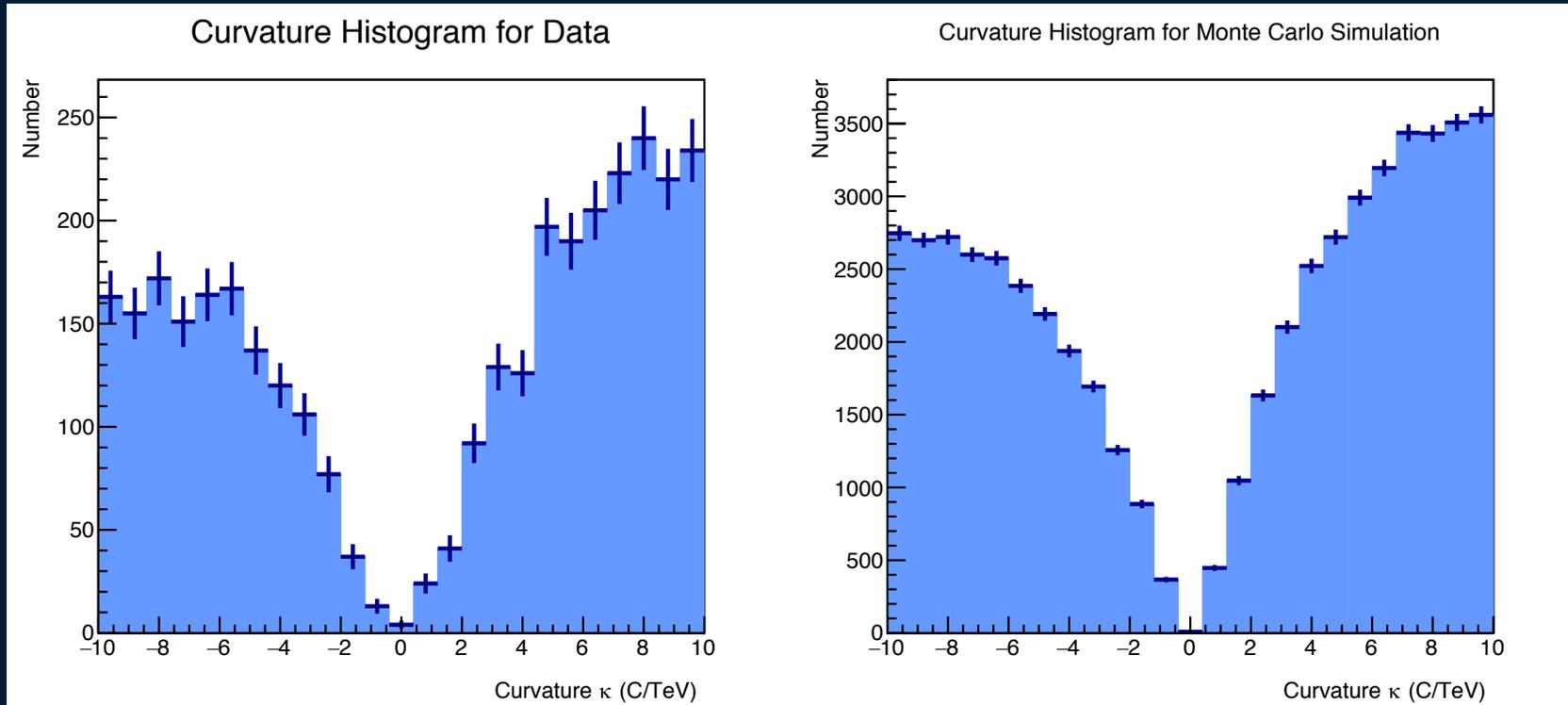
# PROCEDURE

## 1. Calculate and make histograms of $p_T$



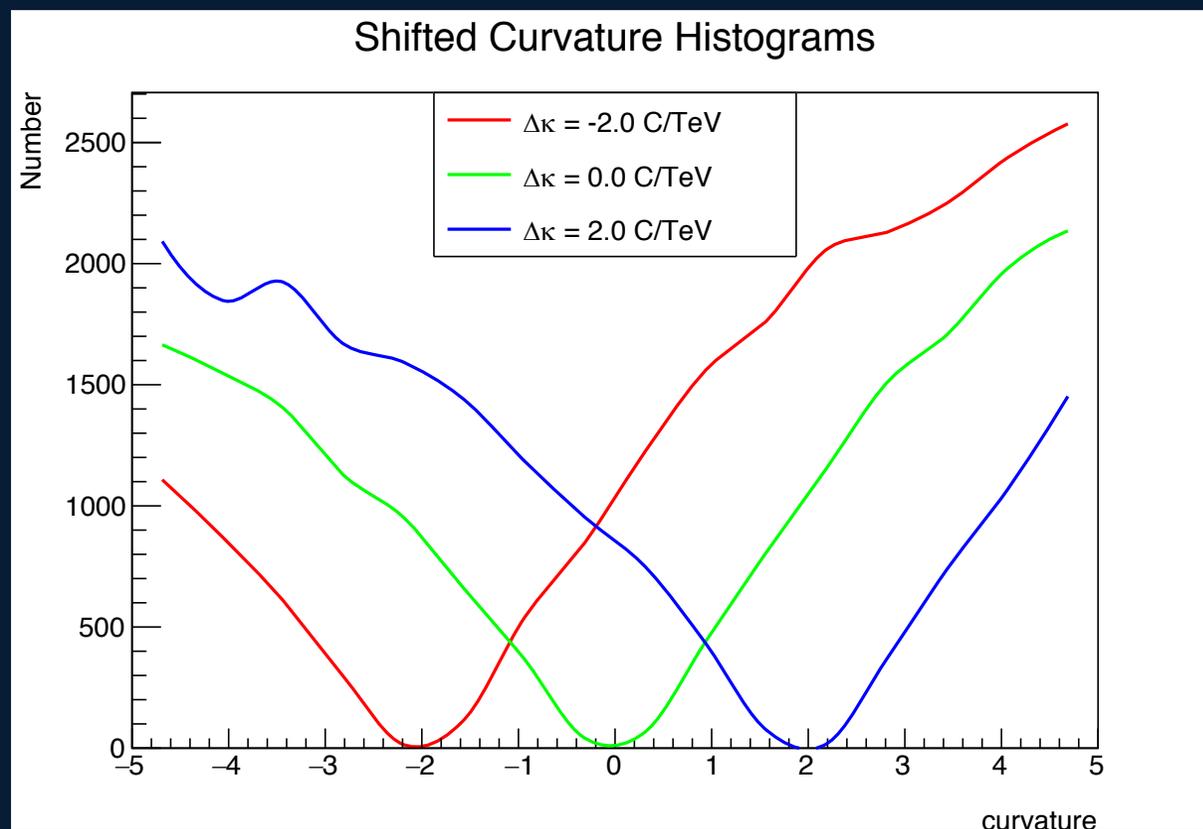
# PROCEDURE

## 2. Calculate and make histograms of $\kappa$

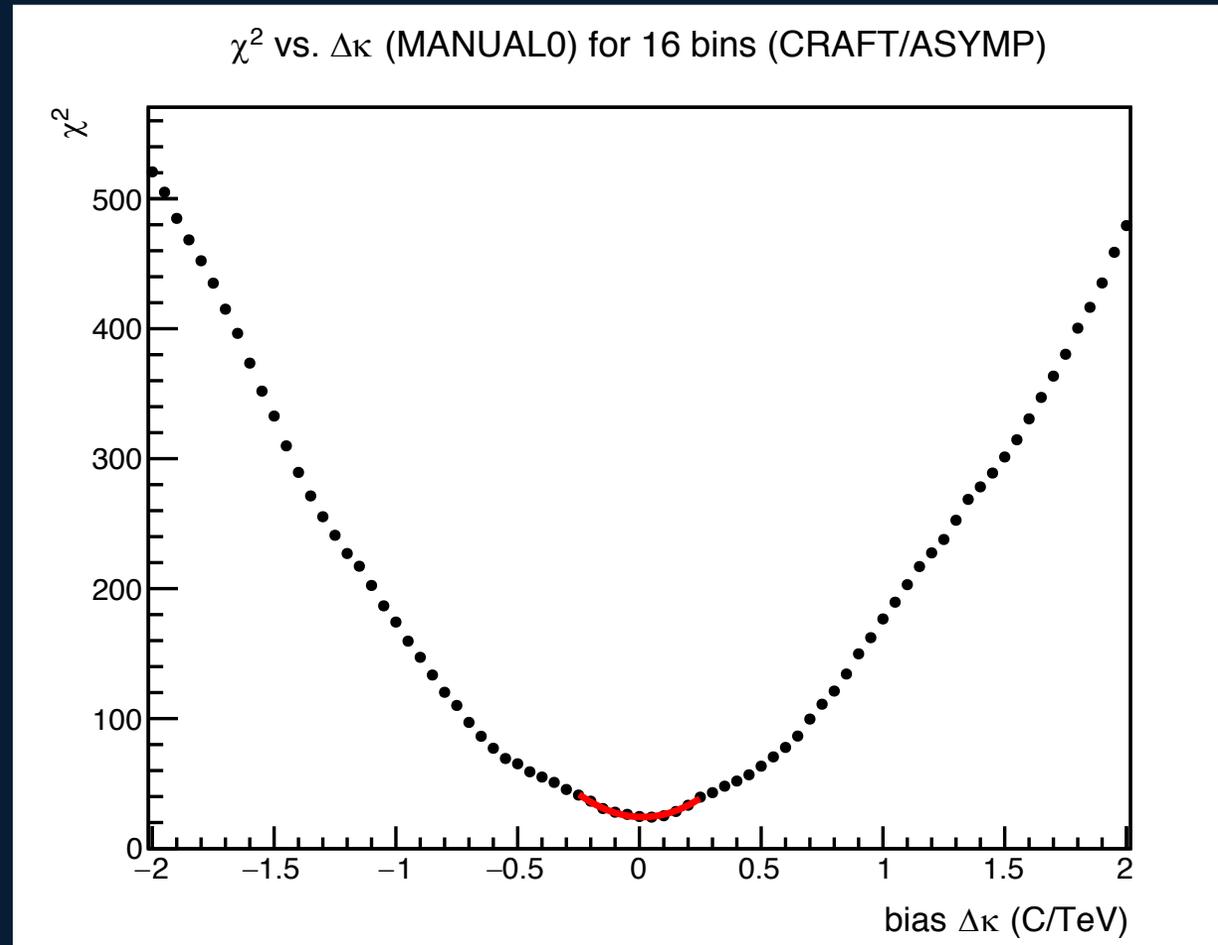


# PROCEDURE

## 3. Apply shift to $\kappa$



# COMPARING DATA AND SIMULATION



# COMPOSITENESS

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- Theory that all elementary particles are actually made up of the same fundamental building blocks, which are called **preons**
- If true, would observe compositeness effects at some energy scale  $\Lambda$ 
  - So far, not observed up to 9.5 TeV in one model of compositeness, and 13.1 TeV in another model

*“The finder of a new elementary particle used to be rewarded by a Nobel Prize, but such a discovery now ought to be punished by a 10,000 dollar fine.”*

*--Willis Lamb*

# QUARK COMPOSITENESS

- At some energy scale  $\Lambda$ , can essentially “break apart” quarks without a Z intermediate and get some fraction of DY events and some fraction of CI events
- However, if  $\Lambda$  is infinite (i.e. it takes infinite energy to break the preons apart) that means the quarks are effectively the smallest indivisible particle
  - In this case, all of the events would be DY

$$\frac{d\sigma^{\text{CI/DY}}}{dM_{\mu\mu}} = \frac{d\sigma^{\text{DY}}}{dM_{\mu\mu}} - \eta_{ij} \frac{I}{\Lambda^2} + \eta_{ij}^2 \frac{C}{\Lambda^4},$$

Additional terms for CI events

